

PHYSICS A2

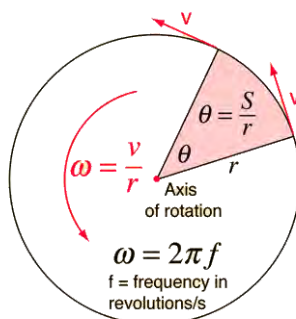
CHAPTER 12 : MOTION IN A CIRCLE

12.1 Kinematics of uniform circular motion

- define the radian and express angular displacement in radians
- understand and use the concept of angular speed

TERMS	DEFINITION/ FORMULA
Angular displacement	<ul style="list-style-type: none"> - Change in angle of a body - as it rotates around a circle
Radian	<ul style="list-style-type: none"> - Angle subtended at the centre of a circle - by an arc of length equal to the radius of the circle
Angular velocity	<ul style="list-style-type: none"> - Rate of change of angular displacement - swept out by radius
Angular speed	<ol style="list-style-type: none"> 1. string: <ul style="list-style-type: none"> - Rate of change of angle - by the string 2. ball : <ul style="list-style-type: none"> - Change in angular displacement - per unit time

- Relationship between v , r and ω



- ω : angular velocity / angular frequency

- Kinetic theory equation :

$$\overline{u^2} = \frac{1}{3} \overline{c^2}$$

$$Pressure = \frac{Force}{Area}$$

$$Force = \frac{Nmc^2}{3l^2}$$

$$P = \frac{Nmc^2}{3l^3}$$

$$P = \frac{Nmc^2}{3V}$$

$$PV = \frac{Nmc^2}{3}$$

- Kinetic energy
 - KE is directly proportional to T

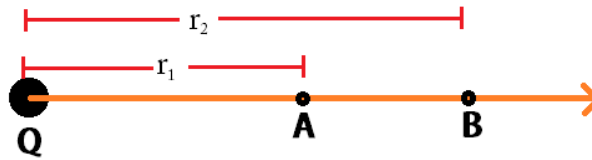
$$KE = \frac{1}{2} Nmc^2$$

$$PV = \frac{Nmc^2}{3} \quad PV = nRT$$

$$Nmc^2 = 3nRT$$

$$KE = \frac{3}{2} nRT$$

- Potential difference



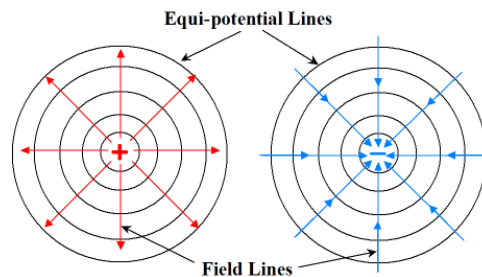
$$V_{AB} = \frac{Q}{4\pi\epsilon_0(r_2 - r_1)}$$

$$V_{AB} = \frac{Q}{4\pi\epsilon_0} \left(\frac{1}{r_2} - \frac{1}{r_1} \right)$$

- V_{AB} is equal to the gain in electrical potential energy if Q is positive and loss if Q is negative

- Equipotential line

- Equipotential surface: a surface where the electric potential is constant
- Equipotential lines are drawn such that potential is constant between intervals
- Potential gradient = 0, $E = 0 \rightarrow$ no work is done when a charge moved along the surface/ line
- Electric field lines must meet equipotential surface at right angles



- Electric potential energy

$$E = Fs$$

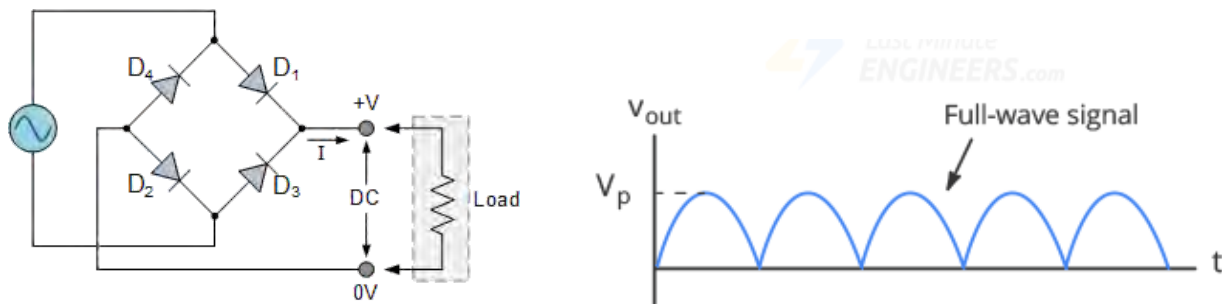
$$F = \frac{Qq}{4\pi\epsilon_0 r^2}$$

$$E = \frac{Qqr}{4\pi\epsilon_0 r^2}$$

$$E = \frac{Qq}{4\pi\epsilon_0 r}$$

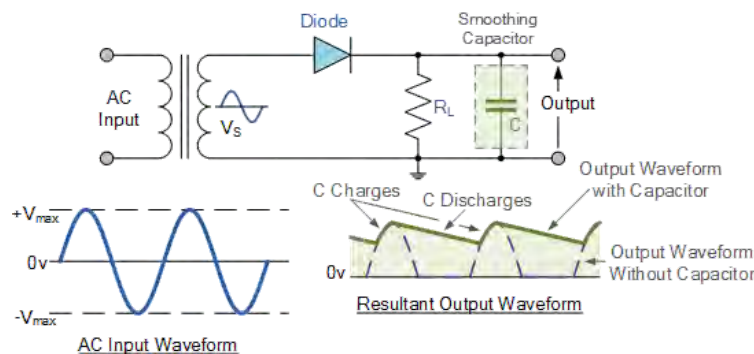
Full-wave rectification

- Use 4 diodes (bridge rectifier)

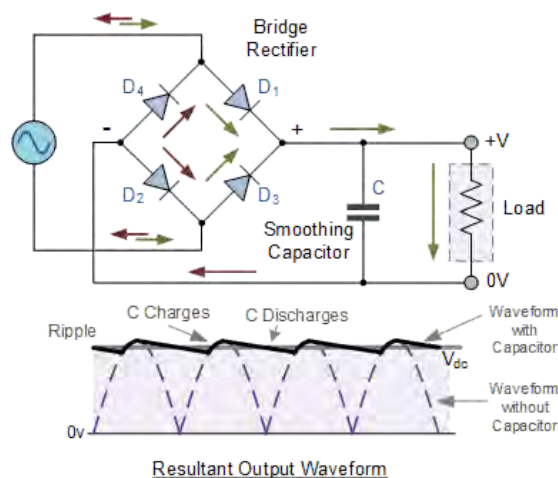


Smoothing

- Use a capacitor to reduce amount of ripple
- The capacitor charges and maintains the voltage as a.c. voltage rises, (first half of the wave)
- As the wave slopes downward, the capacitor begins to discharge in order to maintain the voltage
- **Half wave :**



- **Full wave :**



constant

- There is negligible absorption of this radiated power between the star and the Earth → energy radiated = energy received

Radiant flux intensity/ density (F)

- Radiant flux intensity = Luminosity/ surface area
- Unit : Wm^{-2}

$$F = \frac{L}{A}$$

$$F = \frac{L}{4\pi d^2}$$

Standard candles

- Source of light that has a known luminosity, without having to know its distance
- Example : Cepheid variable stars, Type Ia supernovae
- If we know the luminosity, we can estimate its distance from how bright it appears from Earth

25.2 Stellar radii

- recall and use Wien's displacement law $\lambda_{\text{max}} \propto 1/T$ to estimate the peak surface temperature of a star
- use the Stefan–Boltzmann law $L = 4\pi\sigma R^2 T^4$
- use Wien's displacement law and the Stefan–Boltzmann law to estimate the radius of a star

TERMS	DEFINITION/ FORMULA
Black body	- An idealised object that absorbs all incident electromagnetic radiation
Wien's displacement law	- The black-body radiation curve for different temperatures will peak at different wavelengths that are inversely proportional to the temperature
Stefan-Boltzmann law	- The total energy radiated per unit time per unit surface area of a black body is proportional to the fourth power of its absolute temperature